

# 1 Mantle to surface degassing of carbon- and sulphur-rich alkaline magma at El Hierro, Canary 2 Islands

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## 11 Abstract

12 Basaltic volcanoes transfer volatiles from the mantle to the surface of the Earth. The  
13 quantification of deep volatile fluxes relies heavily on estimates of the volatile content of primitive  
14 magmas, the best archive of which is provided by melt inclusions. Available data from volcanoes  
15 producing mafic alkaline lavas in a range of tectonic settings suggest high volatile fluxes, but  
16 information remains sparse, particularly for intraplate ocean islands. Here we present measurements  
17 of volatile and trace element concentrations, as well as sulphur speciation, in olivine-hosted melt  
18 inclusions and matrix glasses from quenched basanite lava balloon samples from the 2011–2012  
19 submarine eruption at El Hierro, Canary Islands. The results reveal remarkably high concentrations of  
20 dissolved volatiles and incompatible trace elements in this magma, with ~80 ppm Nb and up to 3420  
21 ppm CO<sub>2</sub>, 3.0 wt% H<sub>2</sub>O and 5080 ppm S. Reconstructed primitive CO<sub>2</sub> contents, considering CO<sub>2</sub>/Nb  
22 systematics and possible CO<sub>2</sub> sequestration in shrinkage bubbles, reach weight percent levels,  
23 indicating that carbon is a major constituent of Canary Island magmas at depth and that exsolution of  
24 a CO<sub>2</sub>-rich fluid begins in the mantle at pressures in excess of 1 GPa. Correlations between sulphur  
25 concentration, sulphur speciation and water content suggest strong reduction of an initially oxidised  
26 mantle magma, likely controlled by coupled H<sub>2</sub>O and S degassing. This late-stage redox change may  
27 have triggered sulphide saturation, recorded by globular sulphide inclusions in clinopyroxene and  
28 ulvöspinel. The El Hierro basanite thus had a particularly high volatile-carrying capacity and released a  
29 minimum of 1.3–2.1 Tg CO<sub>2</sub> and 1.8–2.9 Tg S to the environment, causing substantial stress on the  
30 local submarine ecosystem. These results highlight the important contribution of alkaline ocean island  
31 volcanoes, such as the Canary Islands, to volatile fluxes from the mantle.

## 32 Keywords

33 Melt inclusion; carbon dioxide; sulfur; sulfide; degassing; Canary Islands

## 34 1. Introduction

35 Magma degassing of volatile compounds (CO<sub>2</sub>, H<sub>2</sub>O, S, F, Cl) is central to numerous Earth  
36 science problems, ranging from explosive volcanic eruptions, to ore-forming processes, and  
37 atmospheric chemistry and climate change. Deep CO<sub>2</sub> emissions from volcanoes, for example, are a  
38 poorly constrained, yet crucial input for modelling the long-term carbon cycle (Kelemen and Manning,  
39 2015). Sulphur, in turn, with its multiple valences in silicate melts and volcanic gases, participates in  
40 key redox reactions in the mantle and atmosphere (Wallace and Edmonds, 2011). Quantifying fluxes  
41 of deep volatiles towards the surface requires estimates of their primary concentrations in basaltic  
42 magmas (including alkaline varieties) erupted at mid-ocean ridges, subduction zones and intraplate  
43 hotspots. During partial melting of the mantle, volatiles behave like other incompatible elements and  
44 are strongly concentrated in the melt phase (Saal et al., 2002). However, because the solubility of  
45 volatiles in silicate melts is primarily a function of pressure, erupted magmas typically do not preserve  
46 primary melt volatile contents — especially CO<sub>2</sub>, H<sub>2</sub>O and S — due to strong degassing upon ascent-  
47 driven decompression. To circumvent this problem, researchers turn to melt inclusions, which  
48 represent droplets of silicate melt trapped during crystal growth at depth. Protected by their mineral  
49 host acting as a pressure vessel, melt inclusions can retain volatiles during decompression and,  
50 assuming that any post-entrapment modification of their composition can be assessed, provide unique  
51 information on the degassing histories and original concentrations of volatiles in magmas (Métrich and  
52 Wallace, 2008).

53 Alkaline volcanism, in a range of tectonic settings, produces SiO<sub>2</sub>-undersaturated magmas  
54 typically enriched in incompatible trace elements, reflecting low degrees of partial melting of possibly  
55 enriched mantle sources (e.g., McKenzie and O'Nions, 1991). Available melt inclusion data indicate  
56 that mafic alkaline magmas are volatile-rich (Hudgins et al., 2015; Oppenheimer et al., 2011; Spilliaert  
57 et al., 2006) and oxidised (Jugo et al., 2010; Métrich et al., 2009; Métrich and Clocchiatti, 1996;  
58 Moussallam et al., 2014), consistent with experimental evidence (e.g., Dixon, 1997; Lesne et al., 2011;  
59 Shishkina et al., 2014). This suggests that alkaline volcanoes, despite relatively modest magma output  
60 rates, may contribute significantly to deep volatile fluxes (e.g., Oppenheimer et al., 2011), but data  
61 remain sparse, particularly in intraplate oceanic settings.

62 The Canary Islands are a voluminous group of intraplate ocean island volcanoes that  
63 dominantly erupt alkaline magmas derived from a heterogeneous, deep ( $\geq 60$ –100 km) mantle source  
64 (e.g., Day et al., 2010; Gurenko et al., 2006; Sigmarsson et al., 1998). Information pertaining to the  
65 volatile budget of Canary Island magmas is scarce. Wallace (1998) reported CO<sub>2</sub> and H<sub>2</sub>O  
66 concentrations reaching 1430 ppm and 1.4 wt%, respectively, in melt inclusions from basaltic  
67 hyaloclastites from the submarine flanks of Gran Canaria. Gurenko and Schmincke (2000) studied

68 similar samples and observed high S concentrations reaching 5800 ppm, of which up to 95% was  
69 found to be dissolved as sulphate. Abundant data on fluid inclusions exist, however, and attest to the  
70 presence of pervasive CO<sub>2</sub>-rich fluids in the mantle beneath the Canaries (e.g., Hansteen et al., 1991;  
71 Hansteen et al., 1998).

72 From October 2011 to March 2012, a submarine eruption occurred about 2 km south of El  
73 Hierro, the westernmost and youngest island of the archipelago (Fig. 1) (see Carracedo et al., 2015 for  
74 a review). The eruption produced peculiar lava balloons (e.g., Kelly et al., 2014) that were collected  
75 while floating and degassing at the sea surface above the vent. Petrological and geophysical data  
76 indicate that mantle-derived basanite magma intruded the lower crust, where it migrated  
77 subhorizontally for ~15 km, and then rapidly transited to the surface, all within a few months (Longpré  
78 et al., 2014; Martí et al., 2013; Meletlidis et al., 2015).

79 In this paper, we present volatile and trace element concentrations, as well as sulphur  
80 speciation measurements, in olivine-hosted melt inclusions and their host matrix glasses from the El  
81 Hierro lava balloons. This represents the first comprehensive effort to constrain the degassing  
82 behaviour and volatile fluxes at Canary Island volcanoes. The results reveal the strong degassing and  
83 associated redox change of the initially volatile-rich and oxidised 2011–2012 magma as it ascended  
84 from the mantle to produce a Strombolian-type submarine eruption. Our findings also add to a growing  
85 body of evidence of high volatile outputs at alkaline volcanoes.

## 86 **2. Samples and Methods**

87 The quenched lava balloon fragments studied here were erupted at 100–300 m water depth  
88 and collected on 31 October, 27 November, and 6 December 2011, as well as 28 January 2012. We  
89 also analysed the matrix glass composition of the essentially aphyric basanitic crust of a bicoloured  
90 sample erupted on 15 October 2011, representing the minor, earliest eruptive products (e.g., Troll et  
91 al., 2012). Detailed petrologic investigations of the lava balloon samples were previously presented by  
92 Martí et al. (2013), Longpré et al. (2014), and Meletlidis et al. (2015). In brief, the lava contains  
93 phenocrysts of olivine, clinopyroxene, ulvöspinel and rare ilmenite that are set in a glassy matrix  
94 containing microlites of the same minerals, in addition to minor plagioclase. Phenocrysts host  
95 inclusions of Fe-Ti oxides, glass (melt inclusions) and CO<sub>2</sub>-rich fluid. Notably, large (up to 200 µm  
96 across) globular sulphides also frequently occur as inclusions in clinopyroxene and ulvöspinel, often  
97 near or on the rims of crystals (Fig. 2c), but not in olivine nor as a free phase in matrix glass. This work  
98 focuses on melt inclusions (20–150 µm) in olivine, virtually all of which feature vapour bubbles  
99 occupying  $5 \pm 2$  % of the inclusion volume (Fig. 2). Some melt inclusions also contain microlites, most  
100 commonly Fe-Ti oxide, tiny ( $0.1 \pm 0.1$  vol%, i.e.  $\leq 5$  µm) sulphide spheres and clinopyroxene (Fig. 2).

101 We generally targeted glassy melt inclusions lacking evidence of possible leakage and microlites  
102 larger than the vapour bubbles. However, in some cases we could not avoid inclusions with smaller  
103 microlites.

104 CO<sub>2</sub>, H<sub>2</sub>O, S, F and Cl concentrations in 29 melt inclusions and 16 matrix glass chips were  
105 measured by secondary ion mass spectrometry (SIMS) on a CAMECA IMS 1280 instrument at Woods  
106 Hole Oceanographic Institution. A high mass resolution (5000–6000 mass/ $\Delta$ mass) configuration  
107 ensured complete separation of the <sup>12</sup>C and <sup>24</sup>Mg<sup>2+</sup> peaks. Sample preparation protocols and  
108 analytical procedures are outlined by Helo et al. (2011). Ten standard glasses, of basaltic and basaltic  
109 andesite compositions, were used to establish calibration curves for <sup>12</sup>C/<sup>30</sup>Si, <sup>16</sup>O<sup>1</sup>H/<sup>30</sup>Si, <sup>19</sup>F/<sup>30</sup>Si,  
110 <sup>32</sup>S/<sup>30</sup>Si and <sup>35</sup>Cl/<sup>30</sup>Si versus the respective volatile component (Fig. S1, Table S1). The CO<sub>2</sub> content of  
111 all standard glasses was determined by Fourier transform infrared (FTIR) spectroscopy. The standard  
112 error on the slope of the calibration curves is 2.3% for CO<sub>2</sub> and 3.7% or better for H<sub>2</sub>O, F, S and Cl.  
113 Repeated analyses (n=9) of our internal standard P1326-2, a Juan de Fuca Ridge basaltic glass,  
114 during our analytical session yielded 340 ± 24 (1 $\sigma$ ) ppm CO<sub>2</sub>, 0.29 ± 0.09 wt.% H<sub>2</sub>O, 206 ± 6 ppm F,  
115 1296 ± 26 ppm S, and 187 ± 5 ppm Cl (Fig. S2). This agrees well with values obtained by Helo et al.  
116 (2011). One outlier with 0.45 wt.% H<sub>2</sub>O affects the standard deviation for water. Spot analyses of  
117 olivine adjacent to melt inclusions gave 40 ± 15 ppm CO<sub>2</sub>, 0.06 ± 0.00 wt.% H<sub>2</sub>O, 21 ± 2 ppm F, 1 ± 0  
118 ppm S, and 2 ± 0 ppm Cl, demonstrating low analytical backgrounds. Further details on SIMS data  
119 quality and error propagation are provided in the Supplementary Materials.

120 After SIMS analysis, major element compositions of melt inclusions, glasses and olivine were  
121 acquired using a JEOL 8900 electron microprobe (EMPA) at McGill University and are given by  
122 Longpré et al. (2014). We used the same instrument and the method of Wallace and Carmichael  
123 (1994) to measure sulphur speciation in a subset of 17 melt inclusions and 13 matrix glass chips.  
124 Scans of the SK $\alpha$  wavelength were done from 171.517 to 172.478 mm  $\lambda$  with a step size of 3  $\mu$ m and  
125 a dwell time of 0.55 s per step (total 176 s), using a 15  $\mu$ m beam, a beam current of 20 nA, and an  
126 acceleration voltage of 15 kV. Peak shifts of unknowns were measured relative to a sphalerite (ZnS)  
127 standard. Anhydrite (CaSO<sub>4</sub>) and barite (BaSO<sub>4</sub>) standards were used to measure the maximum shift.  
128 In an attempt to quantify possible reduction or oxidation of sulphur upon exposure to the electron  
129 beam (Jugo et al., 2010; Klimm et al., 2012; Métrich and Clocchiatti, 1996; Rowe et al., 2007), we  
130 performed a series of tests using synthetic and natural glasses, the sulphur speciation of which had  
131 previously been measured by XANES and other methods (e.g., Fiege et al., 2015). These tests, fully  
132 presented in the Supplementary Materials (Figs. S3–S4, Table S2), suggest that (1) our EMPA-based  
133 method may significantly underestimate S<sup>6+</sup>/ $\Sigma$ S ratios for oxidised samples (here S<sup>6+</sup>/ $\Sigma$ S ≥ 0.5) (e.g.,  
134 Jugo et al., 2010), (2) beam-induced oxidation of sulphur is negligible in our glasses (cf. Klimm et al.,

135 2012; Rowe et al., 2007), and (3) reproducibility of  $S^{6+}/\Sigma S$  measurements for sulphur-rich glasses is  
136 within 10%, which we assign as a  $1\sigma$  error for unknowns.

137 Trace element concentrations were obtained by laser ablation inductively coupled plasma  
138 mass spectrometry (LA-ICP-MS) at the University of Bremen (matrix glasses), with a New Wave  
139 UP193-SS laser coupled to a Thermo Finnigan Element2 mass spectrometer, and Lamont-Doherty  
140 Earth Observatory (15 melt inclusions and matrix glasses), with a New Wave UP193-FX laser coupled  
141 to a Thermo VG PQ ExCell mass spectrometer. Calibration curves were obtained using international  
142 basaltic glass standards, and repeated in-run analyses of secondary glass standards indicate that our  
143 data are accurate to better than 10% for most elements (see Supplementary Materials for further  
144 details).

145 Melt inclusion compositions were corrected, where applicable, for post-entrapment  
146 crystallisation of the host mineral on the inclusion walls and subsequent diffusive re-equilibration using  
147 the method of Danyushevsky et al. (2000) (see details in Supplementary Materials). On average, the  
148 correction procedure involved addition of 6% olivine. In the text and figures we present the corrected  
149 data, while both the raw and corrected data are listed in full in the Supplementary Materials. As  
150 discussed in sections 3.3 and 3.5, we also explored the effects of shrinkage bubble and sulphide bleb  
151 formation on the respective  $CO_2$  and S budgets of the melt inclusions. These corrections, shown to be  
152 potentially important for  $CO_2$  but relatively minor for S, are presented in the Supplementary Materials  
153 but were not applied to data reported in the text and figures.

### 154 **3. Results and Discussion**

#### 155 *3.1. Petrogenetic relationship between melt inclusions, matrix glasses and the bulk lava*

156 Before discussion of volatile element systematics, it is useful to assess the petrogenetic  
157 relationship between the melt inclusions, the matrix glasses and the bulk lava. Olivine-hosted melt  
158 inclusions show a strong compositional affinity with bulk lava balloons, both in terms of major elements  
159 and elevated concentrations of incompatible trace elements (Fig. 3) (Longpré et al., 2014; Martí et al.,  
160 2013). In comparison, matrix glasses have lower MgO and even more enriched trace element  
161 concentrations. In the context of other mafic lavas from the Canaries, the 2011–2012 melt inclusions  
162 and bulk rocks appear typical for the western islands (El Hierro and La Palma), whereas lavas from  
163 the central islands (La Gomera, Tenerife and Gran Canaria) show lower incompatible trace element  
164 concentrations (Figs. 1, 3b) (Day et al., 2010; Gurenko et al., 2006). In comparison to glasses from  
165 Loihi and Kilauea volcanoes, Hawaii (Dixon and Clague, 2001; Sides et al., 2014), the El Hierro  
166 glasses are markedly more enriched.

167 The trace element data clearly indicate the olivine-hosted melt inclusions to be co-genetic to  
168 the 2011–2012 magma. In detail, the weak positive correlation of trace element enrichment proxies,  
169 such as the Ce/Y ratio, and trace element concentration (cf. Neave et al., 2014) suggests that the  
170 olivine-hosted melt inclusions represent variably enriched, but overall closely-related melts that mixed  
171 to produce the 2011–2012 bulk magma, which evolved via crystallisation of olivine, clinopyroxene and  
172 ulvöspinel to produce a residual melt — the matrix glass. Therefore, a melt more evolved than the  
173 matrix glasses likely did not participate in magma mixing below El Hierro in 2011–2012, and the origin  
174 of evolved (up to basaltic trachyandesite) ulvöspinel-hosted melt inclusions (cf. Longpré et al., 2014)  
175 will be re-evaluated elsewhere. We concur with Meletlidis et al. (2015) that data are consistent with a  
176 single batch of magma, with mixing of similar basanitic melts and entrainment of antecrysts producing  
177 zoned crystals and crystal cumulus in the magma body yielding an increase in crystal (and MgO)  
178 content with time in the bulk lava. The diffusion chronometry results of Longpré et al. (2014) record  
179 mobilisation of this magma batch shortly prior to and during eruption.

### 180 3.2. *A volatile-rich magma*

181 Olivine-hosted melt inclusions exhibit a wide range of volatile contents extending to high  
182 concentrations, revealing the volatile-rich nature of the 2011–2012 magma. Carbon dioxide  
183 concentrations in melt inclusions vary from 60 to 3420 ppm, with 50% of values above 1000 ppm CO<sub>2</sub>.  
184 Dissolved water and sulphur contents are positively correlated and show ranges of 0.4–3.0 wt.% and  
185 520–5080 ppm, respectively (Figs. 4–5). Concentrations of 1200–2210 ppm F and 580–1210 ppm Cl  
186 are positively correlated (Fig. 6). To our knowledge, these CO<sub>2</sub> and H<sub>2</sub>O concentrations are the  
187 highest ever measured in melt inclusions from an ocean island volcano (e.g., Hauri, 2002; Métrich and  
188 Wallace, 2008), but are similar to those reported for other alkaline volcanoes such as Etna and Erebus  
189 (Oppenheimer et al., 2011; Spilliaert et al., 2006). Similarly high sulphur contents have been reported  
190 for submarine samples off the island of Gran Canaria (Gurenko and Schmincke, 2000), and so Canary  
191 Island magmas appear to be particularly sulphur-rich. In comparison to melt inclusions, matrix glasses  
192 contain much lower CO<sub>2</sub> (50–490 ppm), H<sub>2</sub>O (0.4–1.0 wt.%) and S (420–1030 ppm) concentrations  
193 (Figs. 4–5), indicating extensive syn-eruptive degassing of these components. In contrast, the highest  
194 F and Cl concentrations in the El Hierro matrix glasses (1810–2140 ppm F and 920–1020 ppm Cl) are  
195 similar to those of olivine-hosted melt inclusions, implying little to no degassing of halogens (Fig. 6).

196 Estimates of undegassed volatile concentrations in basaltic magmas may be derived from  
197 lithophile trace element proxies, such as Nb (or Ba), Ce and Dy, which are thought to behave similarly  
198 to carbon, water and sulphur, respectively, during partial melting and crystallisation (Rosenthal et al.,  
199 2015; Saal et al., 2002). However, the CO<sub>2</sub>/Nb and S/Dy systematics of basaltic melts should reflect  
200 the interplay between melting of C-bearing and S-bearing accessory minerals and melting of silicates

201 (Dasgupta and Hirschmann, 2010; Dasgupta et al., 2007) and so will in part be controlled by the  
202 original abundances of carbonates (or diamond) and sulphide in the source, which will vary in a  
203 heterogeneous mantle. We thus expect mantle volatile/lithophile trace element ratios to diverge  
204 geographically (e.g., Helo et al., 2011). In this context, the mantle  $\text{CO}_2/\text{Nb}$  and  $\text{S}/\text{Dy}$  ratios of  $240 \pm 50$   
205 and  $225 \pm 50$ , respectively, derived from the Siqueiros melt inclusions representing high-degree, ultra-  
206 depleted melts (Saal et al., 2002) are unlikely to pertain to the low-degree, enriched melts recorded by  
207 the El Hierro melt inclusions. A  $\text{CO}_2/\text{Nb}$  of  $500 \pm 170$  (Rosenthal et al., 2015) and a  $\text{S}/\text{Dy}$  of  $370 \pm 90$   
208 (McDonough and Sun, 1995), thought to be representative of primitive mantle, may be more  
209 appropriate for enriched ocean island basalts. In comparison, measured  $\text{CO}_2/\text{Nb}$  ratios in the 2011–  
210 2012 olivine-hosted melt inclusions are much lower, ranging from 0.7 to 42 (Fig. 7a), which suggests  
211 that undegassed  $\text{CO}_2$  concentrations in El Hierro basanite greatly exceeded observed values. A  
212 simple linear extrapolation to an undegassed  $\text{CO}_2/\text{Nb}$  ratio of  $500 \pm 170$  yields primitive  $\text{CO}_2$  contents  
213 in the range 2.5–5.0 wt%. This implies that  $\geq 85\%$  of the original  $\text{CO}_2$  is missing from the melt  
214 inclusions and that carbon is a major constituent of primitive Canary Island magmas at depth, which is  
215 consistent with predictions of Dixon (1997). In contrast to the  $\text{CO}_2/\text{Nb}$  ratio,  $\text{S}/\text{Dy}$  ranges from 110 to  
216 840, reaching or even exceeding estimated mantle values (McDonough and Sun, 1995), whereas the  
217 upper end of observed  $\text{H}_2\text{O}/\text{Ce}$  ratios (30 to 280) are similar to undegassed values for mid-ocean ridge  
218 and ocean island basalts (155–280) (e.g., Michael, 1995; Saal et al., 2002). This suggests that the  
219 highest  $\text{H}_2\text{O}$  (3.05 wt%) and S (5080 ppm) concentrations in El Hierro melt inclusions may be  
220 minimally degassed (Fig. 7b, c).

### 221 3.3. *Deep $\text{CO}_2$ degassing and pressures of melt inclusion entrapment*

222 The high volatile-carrying capacity of the El Hierro basanite is consistent with experimental  
223 data demonstrating the strong positive effect of alkalinity on  $\text{CO}_2$  solubility (e.g., Dixon, 1997;  
224 Shishkina et al., 2014) (Fig. 4e). For the inferred primitive  $\text{CO}_2$  concentrations reaching weight percent  
225 levels, exsolution of a  $\text{CO}_2$ -rich fluid begins at pressures in excess of 1 GPa, in agreement with fluid  
226 inclusion data reported by Hansteen et al. (1991). For the observed  $\text{CO}_2$  concentrations, however,  
227 calculated vapour saturation pressures are relatively modest, with 3420 ppm  $\text{CO}_2$  corresponding to  
228 260–630 MPa depending on model (Iacono-Marziano et al., 2012; Newman and Lowenstern, 2002;  
229 Papale et al., 2006; Shishkina et al., 2014) and melt composition assumptions, particularly  $\text{Fe}^{3+}/\Sigma\text{Fe}$   
230 ratios (Fig. 4d). The model of Shishkina et al. (2014) is particularly well calibrated for basanitic  
231 compositions and yields an overall saturation pressure range of 8–280 MPa for olivine-hosted melt  
232 inclusions. However, for consistency with subsequent degassing modelling, we adopt the model of  
233 Iacono-Marziano et al. (2012), which is implemented into the D-Compress software (Burgisser et al.,  
234 2015). This gives a similar range of pressures of 7 to 260 MPa, i.e., within the upper 10 km beneath El  
235 Hierro. This is at odds with fluid inclusion and clinopyroxene–melt barometry indicating crystallisation,

236 and hence melt inclusion entrapment, in the uppermost mantle, consistent with pre- and syn-eruptive  
237 seismicity (Fig. 4c) (Longpré et al., 2014). Most notably, CO<sub>2</sub>-rich fluid inclusions record pressures  
238 between 280 and 580 MPa, with most of values above 400 MPa corresponding to the local Moho at  
239 13–15 km depth. This requires that the El Hierro basanite had already begun exsolving a CO<sub>2</sub>-rich  
240 fluid at depths of at least 20 km.

241 Despite the significant uncertainties associated with any barometric method [e.g., fluid  
242 inclusion-derived pressures are dependent upon the chosen equation of state (Hansteen and Klügel,  
243 2008) and clinopyroxene-melt barometry has a standard error of estimate of 150–200 MPa for  
244 calibration data and somewhat higher for global experimental datasets (Putirka, 2008)], we argue that  
245 the complete lack of overlap between pressures derived from CO<sub>2</sub> and H<sub>2</sub>O contents in melt inclusions  
246 and other barometric methods considered here cannot solely be attributed to model errors. We  
247 envisage two possible explanations for this discrepancy:

248 (1) Calculated volatile saturation pressures could underestimate melt inclusion entrapment  
249 pressures because CO<sub>2</sub> originally dissolved in the melt might have later diffused into bubbles, which  
250 are ubiquitous in our samples (Fig. 2). Indeed, recent work has highlighted the possibility that  
251 inclusion-hosted bubbles, which are thought to form in response to post-entrapment processes  
252 affecting the inclusions (see summary by Hartley et al., 2014), may sometimes contain a large fraction  
253 of the CO<sub>2</sub> that was originally dissolved in the silicate melt upon entrapment (Hartley et al., 2014;  
254 Moore et al., 2015). Mean bubble to melt inclusion volume ratios in the El Hierro samples are  $0.05 \pm$   
255  $0.02$ , at the upper end of values reported for olivine-hosted melt inclusions elsewhere (Hartley et al.,  
256 2014; Moore et al., 2015; Neave et al., 2015). A theoretical estimate of how much CO<sub>2</sub> may be  
257 sequestered in the bubble can be obtained using the method of Shaw et al. (2008) based on the ideal  
258 gas law. This calculation, presented in full in the Supplementary Materials, suggests that ~85% of the  
259 originally dissolved CO<sub>2</sub> (but minor H<sub>2</sub>O) may be present in bubbles, yielding reconstructed CO<sub>2</sub>  
260 concentrations of up to 2.3 wt%. This would imply melt inclusion entrapment at high pressures (see  
261 Fig. 4d). As outlined by Neave et al. (2014) and Moore et al. (2015), however, this approach likely  
262 provides an upper bound for the amount of CO<sub>2</sub> sequestered in the bubble, because gas–melt  
263 equilibrium may not have been attained during rapid quenching of the melt inclusions. In addition, we  
264 note that the majority of bubbles investigated by Raman spectroscopy so far actually did not contain  
265 detectable CO<sub>2</sub> (cf. Hartley et al., 2014; Moore et al., 2015; Neave et al., 2014; Neave et al., 2015).

266 (2) Alternatively, melt inclusions entrapped at mantle depths corresponding to pressures  
267 indicated by clinopyroxene–melt and fluid inclusion geobarometry could have lost volatiles during  
268 ascent-driven decompression before sealing off at variable crustal pressures. This implies either  
269 incomplete closure of the melt inclusions at depth or decrepitation of the inclusions followed by healing



270 at different depths (Métrich and Wallace, 2008). Neave et al. (2015) came to the similar conclusion  
271 that melt inclusions from Grímsvötn, Iceland, must have lost CO<sub>2</sub> upon ascent to shallow depths. If this  
272 second possibility is correct, bubbles in El Hierro melt inclusions do not contain significant amounts of  
273 CO<sub>2</sub>, but most of the originally dissolved CO<sub>2</sub> was lost by deep degassing of the primitive basanite  
274 prior to final melt inclusion entrapment. Distinguishing between possibilities (1) and (2) would require  
275 measurements of CO<sub>2</sub> in bubbles by Raman analysis, which is beyond the scope of this study.  
276 Scenario (2), however, is supported by evidence for concurrent H<sub>2</sub>O and S degassing at pressures  
277 ≤300 MPa and is therefore favoured (see section 3.5, and Figs. 5b, 9a).

#### 278 3.4. CO<sub>2</sub> supersaturation in matrix glasses

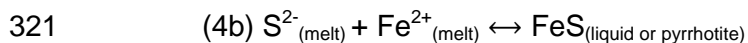
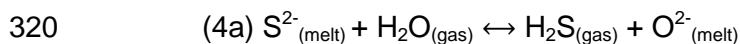
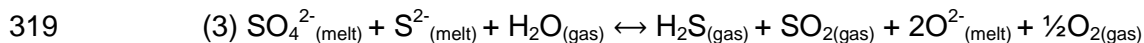
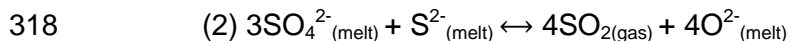
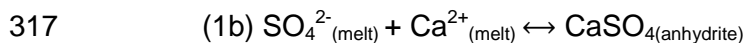
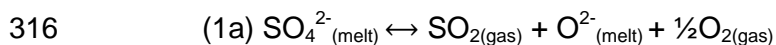
279 A closer look at matrix glasses reveals that the extensive carbon degassing discussed above  
280 was nevertheless incomplete. Figure 4b shows that all individual matrix glass analyses yielded CO<sub>2</sub>  
281 contents significantly in excess of 30 ppm, the expected maximum concentration for an eruption depth  
282 of 300 m below sea level (3 MPa) (Iacono-Marziano et al., 2012). Thus, the magma was  
283 supersaturated in CO<sub>2</sub> when it reached the surface and erupted. Disequilibrium degassing during rapid  
284 magma ascent will result in such CO<sub>2</sub> supersaturation (e.g., Paonita and Martelli, 2006). We note that  
285 matrix glasses from the earliest eruption products are characterised by the highest CO<sub>2</sub>, which likely  
286 reflects higher magma ascent rate, consistent with shorter diffusion timescales obtained at the  
287 outermost rims of olivines from early samples (cf. Longpré et al., 2014).

#### 288 3.5. Sulphur systematics

289 In contrast to most mid-ocean ridge and ocean island basalts (e.g., Wallace and Edmonds,  
290 2011), sulphur concentrations at El Hierro are not correlated with melt iron content and far exceed the  
291 expected sulphur concentration at sulphide saturation (SCSS) of 1520–2650 ppm calculated with the  
292 model of Fortin et al. (2015) (which we select because it accounts for the effect of H<sub>2</sub>O content on the  
293 SCSS) over a wide range of pressures (0.1–300 MPa), temperatures (1100–1200 °C), and water  
294 contents (0–3.5 wt%) (Fig. 5a, b). At first glance this seems at odds with the presence of sulphides in  
295 the samples (Fig. 2). However, the apparently high SCSS in the El Hierro magma is reconciled by the  
296 appreciable amounts of dissolved sulphate (Fig. 5c), which is much more soluble than sulphide in  
297 silicate melts (e.g., Wallace and Edmonds, 2011). Starting with a SCSS of 1520–2020 ppm when all  
298 sulphur is dissolved as S<sup>2-</sup>, the model of Jugo (2009) (his equation 5) predicts an increased SCSS of  
299 6720–8980 ppm for S<sup>6+</sup>/ΣS = 0.7, consistent with our data. Substantial concentrations of dissolved  
300 sulphate appear common in Canary Island magmas (Bryan et al., 2002; Cooper et al., 2015; Gurenko  
301 and Schmincke, 2000; Parat et al., 2011). In a study of submarine glasses from Samoa, Labidi et al.

302 (2015) proposed that high  $S^{6+}/\Sigma S$  ratios in arc and ocean island basalts may reflect the oxidising effect  
303 of melt water content on dissolved sulphur species.

304 Remarkably, both sulphur concentration and the  $S^{6+}/\Sigma S$  ratio in El Hierro melt inclusions are  
305 positively correlated with water content (Fig. 5b, c, d). The preservation of these correlations rules out  
306 any significant influence of post-entrapment diffusion of  $H^+$  (e.g., Bucholz et al., 2013) in controlling the  
307  $H_2O$  content of melt inclusions. To a first order, mixing of volatile-rich and degassed melts, followed by  
308 melt entrapment at variable pressures, could potentially produce the linear relationships observed in  
309 Figure 5. However, simple degassing-induced crystallisation offers a simpler explanation to the broad  
310 negative correlations of  $H_2O$  and S with incompatible elements, such as Na (Fig. 5a inset), K, or Nb.  
311 Therefore, assuming that the melt inclusions are genetically related to each other and to the matrix  
312 glass and that the initial melt composition is best represented by the most volatile-rich inclusion, a  
313 process or processes must have caused a concurrent decrease in S and  $H_2O$  concentrations and the  
314  $S^{6+}/\Sigma S$  ratio in the melt. In Figure 8, we illustrate scenarios 1–4 considered by Marini et al. (2011) (see  
315 their Figure 14) that result in S loss in the melt with expected associated changes in  $S^{6+}/\Sigma S$ :



322 Of the above scenarios, only degassing of oxidised sulphur (1a) or anhydrite precipitation (1b)  
323 appear able to reproduce the S versus  $S^{6+}/\Sigma S$  melt inclusion array. As melt S concentration  
324 decreases, reaction (2) produces a decrease in  $S^{6+}/\Sigma S$  that is too small to account for the observed  
325 variations. Reactions (3), (4a) and (4b) result in increasing  $S^{6+}/\Sigma S$  with decreasing S content, opposite  
326 to our observations. As regards reaction (4b), however, the common presence of single  $\leq 5\ \mu m$   
327 sulphide spheres within the El Hierro melt inclusions (Fig. 2b) deserves further attention. High-  
328 resolution petrographic measurements show that the size of these sulphides is a constant fraction of  
329 the melt inclusion volume, i.e.,  $0.1 \pm 0.1\ vol\%$ , suggesting a common formation mechanism post-  
330 dating melt inclusion entrapment. Danyushevsky et al. (2002) argued that such sulphide blebs may  
331 form in response to a decrease in sulphur solubility brought about by Fe loss from olivine-hosted melt  
332 inclusions, and we envisage a similar scenario for our suite of inclusions. We thus conclude that

333 reaction (4b) did occur in many of our melt inclusions, but had a minor effect on the observed sulphur  
334 systematics, as corroborated by mass balance calculations (see details in the Supplementary  
335 Materials, Fig. S6). In this context and since anhydrite is not observed in the El Hierro samples, ruling  
336 out reaction (1b), it is tempting to conclude that transfer of oxidised sulphur ( $S^{6+}$  dissolved as  $SO_4^{2-}$ )  
337 from the melt to  $SO_2$  gas, in which sulphur occurs as  $S^{4+}$ , coupled to simultaneous  $H_2O$  degassing  
338 generated the trends observed in Figure 5. However, depictions of the effects of reactions 1–4 in  
339 Figure 8 treat the system as a finite pool of reduced and oxidised sulphur species, but sulphur  
340 speciation in silicate melts is sensitive to oxygen fugacity ( $fO_2$ ) and, hence, a redox reaction could  
341 impose a new  $S^{6+}/\Sigma S$  to the system.

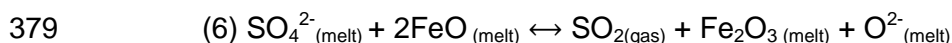
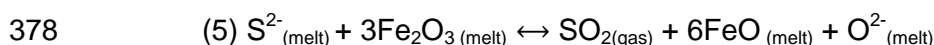
### 342 3.6. *Magma redox*

343 Based on the model of Jugo et al. (2010), we computed  $fO_2$  values using the  $S^{6+}/\Sigma S$  ratio as a  
344 proxy.  $H_2O$ - and S-rich melt inclusions with a high  $S^{6+}/\Sigma S$  ratio should record a higher  $fO_2$  (up to  $\Delta FMQ$   
345 + 1.2, where  $\Delta FMQ$  refers to the  $fO_2$  of the sample relative to the fayalite–magnetite–quartz buffer in  
346 log units) than  $H_2O$ - and S-poor melt inclusions and matrix glasses with low  $S^{6+}/\Sigma S$  (down to  $\Delta FMQ$  +  
347 0.2). In comparison, two-oxide thermobarometry and olivine–liquid equilibrium using crystal rims  
348 suggest more oxidising conditions of  $\Delta FMQ$  + 1.8 and + 2.6, respectively, corresponding to  $Fe^{3+}/\Sigma Fe$   
349 ratios of 0.20 and 0.26 based on the expression of Kress and Carmichael (1991) (Longpré et al.,  
350 2014). We consider it likely that high-end  $fO_2$  estimates obtained with these different methods may  
351 overlap within errors. Blundy and Cashman (2008) computed an average absolute deviation of  $\sim 0.3$   
352 log units for the two-oxide thermobarometer of Ghiorso and Evans (2008), but errors associated with  
353 the other two methods are difficult to quantify. Because of potential photo-reduction of S during  
354 electron microprobe analysis (with the possible counter-effect of post-entrapment sulphide formation),  
355 our  $S^{6+}/\Sigma S$  data may underestimate maximum  $fO_2$  (see section 2 and Supplementary Materials).  
356 However, as sulphur speciation in silicate melts can change extremely rapidly (Métrich et al., 2009),  
357  $S^{6+}/\Sigma S$  measurements may track late-stage  $fO_2$  variability not recorded by mineral–mineral or mineral–  
358 melt thermobarometers. In this context, the melt inclusions appear to capture a reduction of the  
359 magma oxidation state by at least one log unit of  $fO_2$  upon ascent and degassing. Moussallam et al.  
360 (2014) reported a similar situation for basanitic to phonolitic samples from Erebus volcano, Antarctica.  
361 These authors observed a positive correlation of  $Fe^{3+}/\Sigma Fe$  ratios (measured by XANES) in melt  
362 inclusions and calculated vapour saturation pressures, translating into a decrease of  $fO_2$  by more than  
363 2 log units during magma ascent and degassing.

364 To quantify the degassing process and its impact on  $fO_2$  evolution, we conducted a series of  
365 simulations using the D-Compress software (Burgisser et al., 2015). The El Hierro matrix glass was  
366 used as melt composition with initial  $CO_2$  and  $H_2O$  concentrations of 3420 ppm and 3.05 wt%,

367 respectively, at 300 MPa, 1150 °C, and  $fO_2 = FMQ + 1.2$  and  $+1.7$ . Notably, initial melt sulphur  
368 concentration is not an input in D-Compress; it is computed based on implemented solubility laws. The  
369 key results of these degassing models, distilled in Figure 9, are as follows. (1) Coupled S and H<sub>2</sub>O  
370 degassing reproduces the S vs. H<sub>2</sub>O melt inclusion array and is predicted to begin at relatively high  
371 pressure (300 MPa) (cf. Lesne et al., 2011). (2) All realistic open system and closed system degassing  
372 scenarios yield a net decrease in melt  $fO_2$  upon decompression, consistent with our  $S^{6+}/\Sigma S$  data (Fig.  
373 5c). This strongly supports our conclusion that extensive decompression-driven S and H<sub>2</sub>O degassing  
374 of the El Hierro basanite caused a significant reduction of its oxidation state.

375 It remains difficult to pinpoint the specific reactions governing this redox change. Moussallam  
376 et al. (2014) concluded that sulphur degassing is driving the redox evolution of Erebus magmas  
377 according to the following reactions:



380 Based on the observed decrease in  $Fe^{3+}/\Sigma Fe$  ratios at Erebus, these authors favoured reaction  
381 (5). As this reaction proceeds to the right, the melt is reduced via SO<sub>2</sub> degassing and, consequently,  
382 melt  $S^{6+}/\Sigma S$  will re-equilibrate to a new, lower  $fO_2$ . This is consistent with XANES spectra for sulphur  
383 indicating that Erebus melts entrapped at higher pressures are characterised by both higher  $Fe^{3+}/\Sigma Fe$   
384 and  $S^{6+}/\Sigma S$  (Moussallam et al., 2014). Preliminary XANES data for El Hierro samples also show that  
385 melt inclusions have higher  $Fe^{3+}/\Sigma Fe$  ratios than degassed matrix glasses (Y. Moussallam, personal  
386 communication 2015), mirroring our  $S^{6+}/\Sigma S$  data and consistent with an important role for reaction (5).

### 387 3.7. Sulphide saturation

388 The occurrence of large, conspicuous sulphide globules as inclusions near the rims of  
389 clinopyroxene and ulvöspinel (Fig. 2c) appears to record sulphide saturation in the El Hierro magma.  
390 In contrast, the tiny inclusion-hosted sulphides discussed above most likely formed in response to  
391 post-entrapment modification of melt inclusion composition and thus have little to do with bulk melt  
392 sulphide saturation. The virtual absence of the large sulphide inclusions in olivine (inspection of  
393 several hundred olivine grains in the course of this study returned a single sulphide inclusion) and as  
394 free floating blebs in matrix glass may suggest that sulphide saturation occurred late, but that  
395 sulphides ultimately became unstable and were (re)dissolved in the silicate melt or scavenged by a  
396 gas phase prior to eruption. The cause and timing of sulphide saturation in the sulphur-rich El Hierro  
397 basanite are thus intriguing, and we here outline three possibilities. (1) Sulphide saturation was  
398 caused by magnetite crystallisation. Jenner et al. (2010) presented a compelling case for the

399 potentially important role of magnetite crystallisation and associated reduction of the melt in triggering  
400 sulphide saturation in arc-related magmas. If this process was dominant at El Hierro, we would expect  
401 a positive correlation between iron content and the  $S^{6+}/\Sigma S$  ratio in melt inclusions. But these variables  
402 are weakly negatively correlated (Fig. S7), with a drop in FeO observed for matrix glasses only,  
403 suggesting that ulvöspinel crystallisation did not affect the composition of most olivine-hosted melt  
404 inclusions and did not drive the decrease in  $S^{6+}/\Sigma S$  ratio. A subtle change in slope in the S versus FeO  
405 trend at ~1500 ppm S could reflect the late onset of coincidental ulvöspinel crystallisation and sulphide  
406 formation (Fig. 5a). (2) Sulphide saturation was caused by a diffusion-controlled enrichment of sulphur  
407 at the crystal–melt interface (Bacon, 1989). In this case, sulphide saturation would have been reached  
408 only locally, in the vicinity of growing crystals, resulting in the formation and subsequent trapping of  
409 sulphide globules. While this possibility is consistent with the lack of sulphides in the matrix glass, it  
410 calls for an unknown mechanism — although potentially related to differential crystal growth rates —  
411 to produce abundant sulphides in ulvöspinel and clinopyroxene and essentially none in olivine. (3) The  
412 degassing-induced decrease in  $fO_2$  and associated drop in SCSS (e.g., Jugo, 2009) caused sulphide  
413 saturation. As illustrated in Figure 5c, the SCSS is drastically reduced upon decrease of the  $S^{6+}/\Sigma S$   
414 ratio in the melt. In our preferred scenario, the initially oxidised, sulphide-undersaturated El Hierro  
415 magma became sulphide-saturated during degassing as the decrease in SCSS outpaced sulphur loss.

### 416 3.8. *Implications for the mantle source and origin of volatiles*

417 Several lines of evidence, including high  $TiO_2$  contents (Prytulak and Elliott, 2007), multiple  
418 isotopic proxies (Day et al., 2010; Day et al., 2009; Sigmarsson et al., 1998) and olivine compositions  
419 (Gurenko et al., 2009), suggest that the mantle source of El Hierro and other Canary Island lavas  
420 contains significant amounts of recycled oceanic crust, in the form of pyroxenite or eclogite. Day et al.  
421 (2009) estimated the amount of recycled oceanic lithosphere present in the source of El Hierro at  
422 <10%, and Day et al. (2010) presented a partial melting model, using La/Yb and Dy/Yb systematics,  
423 which indicates that 2–6 % melting of an enriched source, beginning in the garnet stability field at  $\geq 110$   
424 km depth and extending into the spinel stability field, can account for the trace element characteristics  
425 of Pleistocene and Holocene El Hierro and La Palma lavas. La/Yb and Dy/Yb ratios for the 2011–2012  
426 samples are  $26.0 \pm 3.0$  and  $3.4 \pm 0.4$ , respectively, overlapping with data reported by Day et al.  
427 (2010), therefore suggesting similar melting conditions. We note, however, that melting of carbonated  
428 peridotite at high pressure may also play a role in the genesis of El Hierro lavas (Dasgupta et al.,  
429 2007).

430 The high pressure — both calculated and inferred — of initial saturation with a  $CO_2$ -rich fluid  
431 and the enrichment of volatile and non-volatile incompatible elements in the El Hierro basanite indicate  
432 that the bulk of volatiles originate from the mantle. Simple batch melting models illustrate the feasibility

433 of generating high initial volatile concentrations by low degrees of partial melting; mismatch by one  
434 order of magnitude of predicted versus observed CO<sub>2</sub> concentrations reinforces our conclusion that  
435 appreciable amounts of CO<sub>2</sub> are missing from melt inclusions (Fig. 10). In addition, low Cl/K ratios  
436 (0.07 ± 0.01) for El Hierro melt inclusions and matrix glasses rule out a significant volatile contribution  
437 from shallow, seawater-influenced components, such as brines, as suggested for some MORB and  
438 Hawaiian glasses (e.g., Kent et al., 1999). Finally, for both El Hierro (this work) and Gran Canaria  
439 (Gurenko and Schmincke, 2000), the most primitive, volatile-rich melt inclusions also show the highest  
440 S<sup>6+</sup>/ΣS ratios, which leads us to propose that the Canary mantle source may be particularly oxidised,  
441 such as inferred for Erebus volcano (Moussallam et al., 2014). This is in agreement with major  
442 element systematics indicating that Canary Island magmas are generally more oxidised than  
443 archetypal ocean island basalts from Hawaii and Iceland (Herzberg and Asimow, 2008).

### 444 3.9. *Eruption style, volatile yields and environmental impact*

445 We propose that initially high volatile contents and intense shallow level degassing led to  
446 submarine Strombolian-type explosive activity at El Hierro. This is consistent with the construction of a  
447 200–300-m-tall submarine cone (Rivera et al., 2013) and strong bubbling at the ocean surface, and it  
448 lends support to an (mildly) explosive origin of lava balloons (Kelly et al., 2014). Based on an eruptive  
449 volume of 329 × 10<sup>6</sup> m<sup>3</sup> (Rivera et al., 2013) converted to ~160–260 × 10<sup>6</sup> m<sup>3</sup> of dense magma  
450 assuming a deposit porosity of 20–50%, a magma crystallinity of 15%, a magma density of 2,900  
451 kg/m<sup>3</sup> and assuming 3200 ppm CO<sub>2</sub> and 4500 ppm S total degassing, we estimate that the El Hierro  
452 eruption released at least 1.3–2.1 Tg CO<sub>2</sub> and 1.8–2.9 Tg S to the environment. We emphasise that  
453 these figures must be regarded as minima, particularly for CO<sub>2</sub>. However, the late formation of  
454 sulphide globules may have provided a sink for some of the sulphur. For comparison, Santana-  
455 Casiano et al. (2013) calculated a CO<sub>2</sub> output of 0.05 Tg/day for the week of 4–9 November 2011  
456 based on pCO<sub>2</sub> values measured in the seawater column in the vicinity of the vent. Extrapolating such  
457 CO<sub>2</sub> discharge over the entire course of the eruption (147 days) yields 7.4 Tg CO<sub>2</sub>. This is likely an  
458 upper limit for the CO<sub>2</sub> yield, as the week of 4–9 November coincided with the most intense bubbling  
459 and degassing at the sea surface. Degassing of the El Hierro basanite caused dramatic decreases in  
460 seawater pH in the vicinity of the vent from a background of 8.0 to a low of 5.1, and anomalously high  
461 concentrations of dissolved reduced sulphur species were also detected in the water column  
462 (Santana-Casiano et al., 2013). On land, detected diffuse CO<sub>2</sub> emissions were comparatively small,  
463 with an estimated 1.1 × 10<sup>-4</sup> Tg CO<sub>2</sub> for the July 2011–March 2012 period (Melián et al., 2014).

464 Our calculated CO<sub>2</sub> yield corresponds to at least 1–7% of the estimated annual flux of 29–154  
465 Tg CO<sub>2</sub> from mid-ocean ridges and ocean islands (Kelemen and Manning, 2015). And while deep  
466 sulphur fluxes from mid-ocean ridges and ocean islands are poorly constrained at present, the S yield

467 of the 2011–2012 eruption equates to 17–28% of Andres and Kasgnoc (1998)'s estimate for the time-  
468 averaged global subaerial emissions of 10.4 Tg/yr S and 23–38% of Shinohara's (2013) 7.7 Tg/yr S  
469 value for subduction zones only. Therefore, despite its modest volume, representing about 1–2% of  
470 the annual magma output at mid-ocean ridge and ocean island volcanoes (Crisp, 1984), the 2011–  
471 2012 El Hierro eruption transferred significant amounts of volatiles to the Earth's surface.

#### 472 **4. Conclusions**

473 In summary, we document some of the highest volatile concentrations ever measured in  
474 silicate melt inclusions from an ocean island volcano. Prior to deep degassing, CO<sub>2</sub> concentrations in  
475 the parental 2011–2012 El Hierro magma likely reached weight percent levels; we envisage carbon to  
476 be a major constituent of primitive Canary Island magmas. The most primitive, least-degassed melt  
477 inclusions are oxidised, and we suggest that this is inherited from the mantle source. Coupled H<sub>2</sub>O  
478 and S degassing beginning at pressures of at least 300 MPa, and continuing up to the surface, caused  
479 a strong late-stage reduction of the oxidation state of the basanite by about 1 log unit of  $fO_2$ . In turn,  
480 this change in magma redox reduced sulphur solubility to the point of sulphide saturation, producing  
481 abundant sulphide globules. The large amounts of volatiles released by degassing of the 2011–2012  
482 magma caused significant stress to the local ecosystem, reflected in seawater chemistry during the  
483 eruption. We propose that mafic alkaline volcanism in intraplate ocean island settings contribute large  
484 amounts of volatiles to the Earth surface.

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496

497 **Figure captions**

498 **Figure 1: The Canary Islands.** The red triangle indicates the site of the 2011–2012 submarine  
499 eruption just south of El Hierro. The inset shows the location of the archipelago in the eastern Atlantic  
500 Ocean. Map constructed with GeoMapApp ([www.geomapapp.org](http://www.geomapapp.org)).

501 **Figure 2: Photomicrographs of the El Hierro samples.** **a**, Melt inclusion-bearing olivine crystal  
502 floating in matrix glass. **b**, Zoom on an olivine-hosted melt inclusion containing a bubble and a  
503 sulphide sphere. **c**, Sulphide inclusion occurring near rim of an ulvöspinel crystal.

504 **Figure 3: Incompatible trace element enrichment.** **a**, Ce/Y ratio versus Nb concentration for  
505 individual 2011–2012 olivine-hosted melt inclusions (this work), and average matrix glass (this work)  
506 and bulk rocks (BR) (Martí et al., 2013). Average bulk rock values for mafic lavas (40–50 wt% SiO<sub>2</sub>) for  
507 El Hierro (EH), La Palma (LP), La Gomera (LG), Tenerife (TF) and Gran Canaria (GC) (Day et al.,  
508 2010; Gurenko et al., 2006) are also shown, as well as melt inclusion and matrix glass data for Hawaii  
509 (Dixon and Clague, 2001; Sides et al., 2014) and ocean floor basalt (OFB) glasses worldwide (Jenner  
510 and O'Neill, 2012). All error bars are  $\pm 1\sigma$ . **b**, Primitive mantle (PM)-normalised (McDonough and Sun,  
511 1995) multi-element diagram with data groupings and sources as in (a), except that here the average  
512 is plotted for the El Hierro melt inclusions, Hawaii (split for Loihi and Kilauea volcanoes) and ocean  
513 floor basalts.

514 **Figure 4: High-pressure CO<sub>2</sub> degassing.** **a**, Dissolved CO<sub>2</sub> and H<sub>2</sub>O concentrations in olivine-hosted  
515 melt inclusions (MI-ol) and average matrix glass measured by ion microprobe. Fully propagated error  
516 bars are shown (see Supplementary Information), except when smaller than symbol size. Isobars for  
517 100, 200, 300 and 400 MPa are shown and were calculated with D-Compress (Burgisser et al., 2015)  
518 using the matrix glass composition and a temperature of 1150 °C. Data for Hawaii from Dixon and  
519 Clague (2001), Hauri (2002) and Sides et al. (2014). **b**, Zoom below 1000 ppm CO<sub>2</sub> on individual  
520 matrix glass analyses. October 15 matrix glasses contain more CO<sub>2</sub> than later samples but all are  
521 CO<sub>2</sub>-supersaturated with respect to their eruption pressure of 3 MPa. Isobars for 3, 25, 50 and 100  
522 MPa are shown. **c**, Histogram of depths (converted from pressures, see Longpré et al. (2014))  
523 recorded by matrix glasses, melt inclusions, fluid inclusions (FI), clinopyroxene–melt geobarometry  
524 (Cpx) and pre- and syn-eruptive volcano-tectonic seismicity (EQ) (Longpré et al., 2014; IGN  
525 earthquake catalogue, [www.ign.es](http://www.ign.es)). Vapour saturation pressures for matrix glasses and melt  
526 inclusions were calculated after Iacono-Marziano et al. (2012). **d**, Comparison of CO<sub>2</sub> solubility  
527 models. CO<sub>2</sub> solubility as a function of pressure was calculated at 1150 °C for the 2011–2012 average  
528 matrix glass composition, assuming equilibrium with a pure CO<sub>2</sub> fluid. SiO<sub>2</sub>-corrected (45.06 wt%)  
529 VolatileCalc (Newman and Lowenstern, 2002) curve (1) is independent of melt iron content. However,



530 the model of Papale et al. (2006) requires FeO and Fe<sub>2</sub>O<sub>3</sub> inputs, which exert strong control on CO<sub>2</sub>  
531 solubility. Curves (2) and (3) respectively assuming molar Fe<sup>3+</sup>/ΣFe = 0 (all iron as FeO) and Fe<sup>3+</sup>/ΣFe  
532 = 0.26 (plausible value based on olivine-glass equilibrium (Longpré et al., 2014)) are shown. The  
533 formulation of Shishkina et al. (2014) is weakly dependent on melt iron content and is used here  
534 assuming all iron as FeO (curve 4). Curve (5) shows the model of Iacono-Marziano et al. (2012). Note  
535 that the solubility models are generally not well calibrated above 500 MPa and so extrapolation to  
536 higher pressure is for illustration purposes only. **e**, Effect of composition on CO<sub>2</sub> solubility; CO<sub>2</sub>  
537 solubility (after Iacono-Marziano et al., 2012) as a function of pressure for El Hierro, average Kilauea  
538 and ocean floor basalt glass compositions.

539 **Figure 5: Sulphur systematics.** Sulphur concentration versus total iron (FeO\*) (a), H<sub>2</sub>O (b) and  
540 sulphur speciation (mole fraction of sulphate over total sulphur measured by electron microprobe) (c)  
541 in melt inclusions and matrix glasses. The inset in (a) shows S (same axis scale) versus Na<sub>2</sub>O and the  
542 effect of concurrent degassing and crystallisation. Dotted lines in (b) show the sulphur concentration at  
543 sulphide saturation (SCSS) at S<sup>6+</sup>/ΣS = 0 calculated after Fortin et al. (2015) at temperatures of 1100  
544 (orange), 1150 (red) and 1200 °C (dark red) over a range of pressures (0.1–300 MPa) and H<sub>2</sub>O  
545 contents (0–3.5 wt%) consistent with our data. The SCSS values at H<sub>2</sub>O = 0 wt% (1520–2020 ppm S)  
546 were then extrapolated to higher S<sup>6+</sup>/ΣS in (c) using Equation 5 of Jugo (2009). **d**, H<sub>2</sub>O content versus  
547 sulphur speciation. Data for ocean floor basalts and Hawaii as in Figures 3 and 4.

548 **Figure 6: Behaviour of halogens.** Fluorine and chlorine concentrations are positively correlated in El  
549 Hierro melt inclusions. The high concentration in matrix glass indicates that very little to no degassing  
550 of these elements occurred upon eruption. Data for Hawaii as in Figure 4.

551 **Figure 7: Volatile/lithophile element systematics.** Ratios of volatile and lithophile trace elements of  
552 similar compatibility are plotted against denominator concentrations and are compared to expected  
553 ratios for undegassed mantle melts (red dotted lines: McDonough and Sun, 1995; orange solid lines:  
554 Rosenthal et al., 2015; dark red dashed lines: Saal et al., 2002). **a**, CO<sub>2</sub>/Nb versus Nb; **b**, H<sub>2</sub>O/Ce  
555 versus Ce; and **c**, S/Dy versus Dy. Whereas H<sub>2</sub>O/Ce and S/Dy ratios in El Hierro melt inclusions reach  
556 or exceed expected mantle values, CO<sub>2</sub>/Nb ratios are much lower, suggesting extensive loss of CO<sub>2</sub>.  
557 In all cases, broad negative correlations between volatile/lithophile ratios and incompatible trace  
558 element concentrations indicate the combined effect of degassing and crystallisation (labelled arrows  
559 in b), whereas displacement of the El Hierro data to the right of the diagrams, with respect to ocean  
560 floor basalts and Hawaii (data as in Figure 3), is due to lower degrees of melting of an enriched  
561 source.

562 **Figure 8: Sulphur redox scenarios.** Sulphur concentration versus speciation, melt inclusion data are  
563 shown by open circles. Curves labelled with numbers correspond to the effects of reactions (1), (2), (3)  
564 and (4), respectively, using initial conditions of 5080 ppm S and  $S^{6+}/\Sigma S = 0.71$ .

565 **Figure 9: Degassing simulations.** **a** and **b** sulphur concentration vs.  $H_2O$  content and  $fO_2$  expressed  
566 as  $\Delta FMQ$ , respectively. Degassing models calculated with D-Compress (Burgisser et al., 2015) are  
567 shown for open (dashed lines) and closed (solid lines) C–S–O–H–Fe systems using the matrix glass  
568 composition and initial conditions of 300 MPa, 1150 °C, 3420 ppm  $CO_2$ , 3.05 wt%  $H_2O$ , and  $fO_2 =$   
569  $FMQ + 1.2$  (grey lines) or  $FMQ + 1.7$  (black lines). **c** and **d** pressure vs.  $fO_2$  during closed and open  
570 system degassing, respectively, with the effect of the system considered, either O–H, S–O–H, C–S–  
571 O–H or C–S–O–H–Fe, shown by curves of different colours. Calculations performed as for **a** and **b**,  
572 but only results for an initial  $fO_2$  of  $FMQ + 1.2$  are shown. For a closed system, only degassing in the  
573 O–H system yields to oxidation of the melt; in all other situations, reduction occurs. For an open  
574 system, degassing in the C–S–O–H–Fe system first produces a slight oxidation, then reduction at  
575 pressure below  $\sim 200$  MPa; in all other cases, reduction occurs. In all panels, open circles show melt  
576 inclusion data, with  $fO_2$  values calculated from  $S^{6+}/\Sigma S$  ratios after Jugo et al. (2010) and saturation  
577 pressures after Iacono-Marziano et al. (2012).

578 **Figure 10: Batch melting constraints on elemental concentrations in the El Hierro basanite.**  
579 Simple batch melting calculations were performed assuming bulk peridotite/melt partition coefficients  
580 ( $D^{peridotite/melt}$ ) of 0.00055 for  $CO_2$ , 0.011 for F, 0.00012 for Ba, 0.079 for Dy and 0.0034 for Nb  
581 (Rosenthal et al., 2015; Workman and Hart, 2005).  $D^{peridotite/melt}$  for S is poorly constrained, but is  
582 thought to be similar to that of rare earth elements (Saal et al., 2002); curves with triangular and  
583 circular symbols assume behaviours similar to those of La ( $D^{peridotite/melt} = 0.01$ ) and Dy ( $D^{peridotite/melt} =$   
584  $0.079$ ), respectively (Workman and Hart, 2005). Solid curves assume a source similar to primitive  
585 mantle (PM) with 600 ppm  $CO_2$ , 300 ppm S, 25 ppm F, 6.6 ppm Ba, 0.674 ppm Dy and 0.75 ppm Nb  
586 (McDonough and Sun, 1995; Rosenthal et al., 2015). Dashed lines are drawn for a depleted MORB  
587 mantle (DMM) source containing 75 ppm  $CO_2$ , 146 ppm S, 16 ppm F, 0.563 ppm Ba, 0.505 ppm Dy  
588 and 0.1485 ppm Nb (Rosenthal et al., 2015; Saal et al., 2002; Workman and Hart, 2005). Digits on  
589 curves refer to melt weight fractions in %. Data for the El Hierro basanite show the highest melt  
590 inclusion values for volatile elements (3420 ppm  $CO_2$ , 5080 ppm S, 2210 ppm F) and average bulk  
591 rock values from Martí et al. (2013) for non-volatile elements (390 ppm Ba, 7.4 ppm Dy, 68 ppm Nb).

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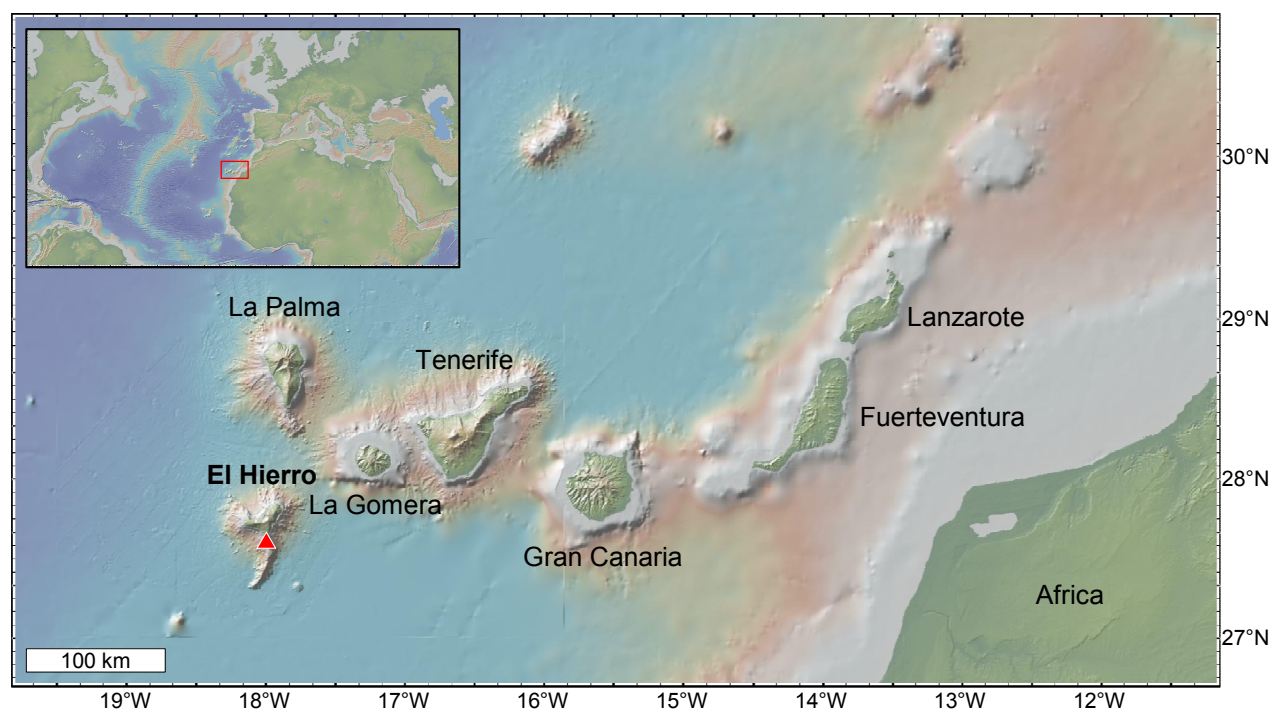
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**Figure**

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**Figure 1**



Figure

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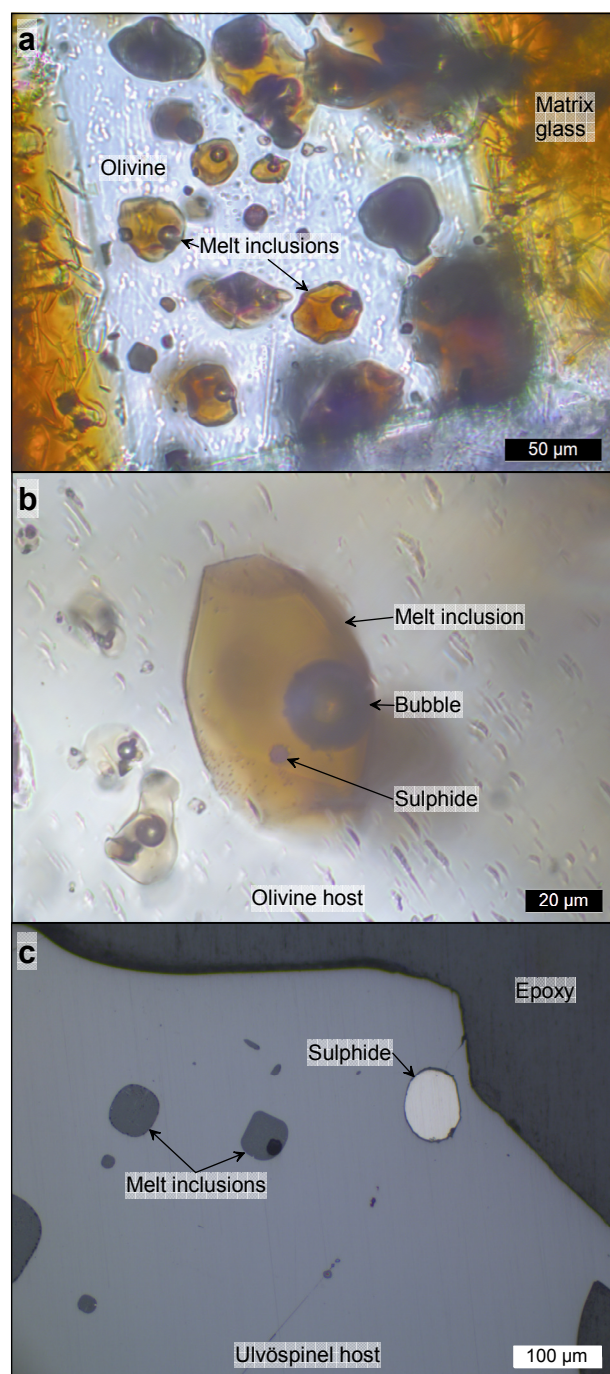
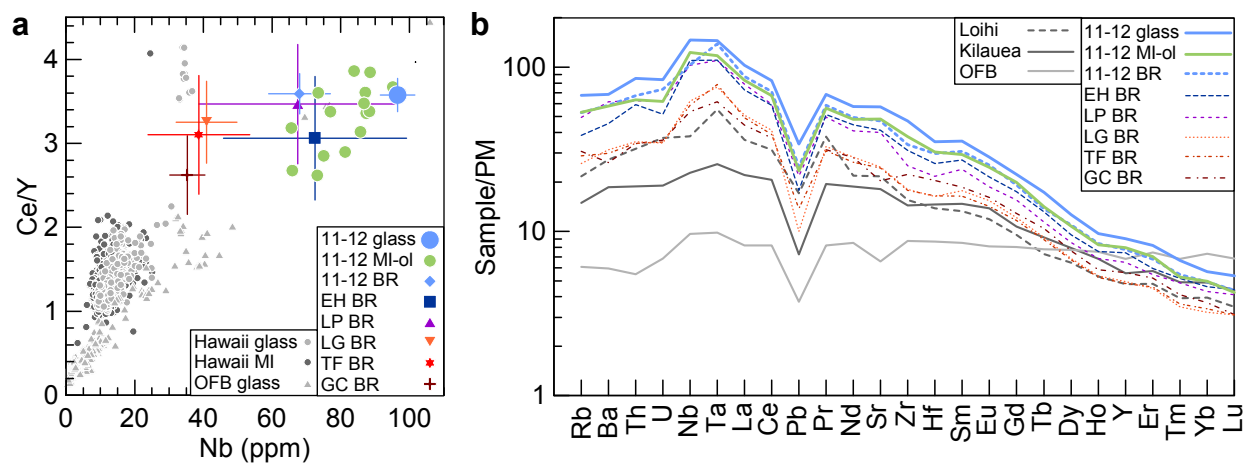


Figure 2

# Figure

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**Figure 3**

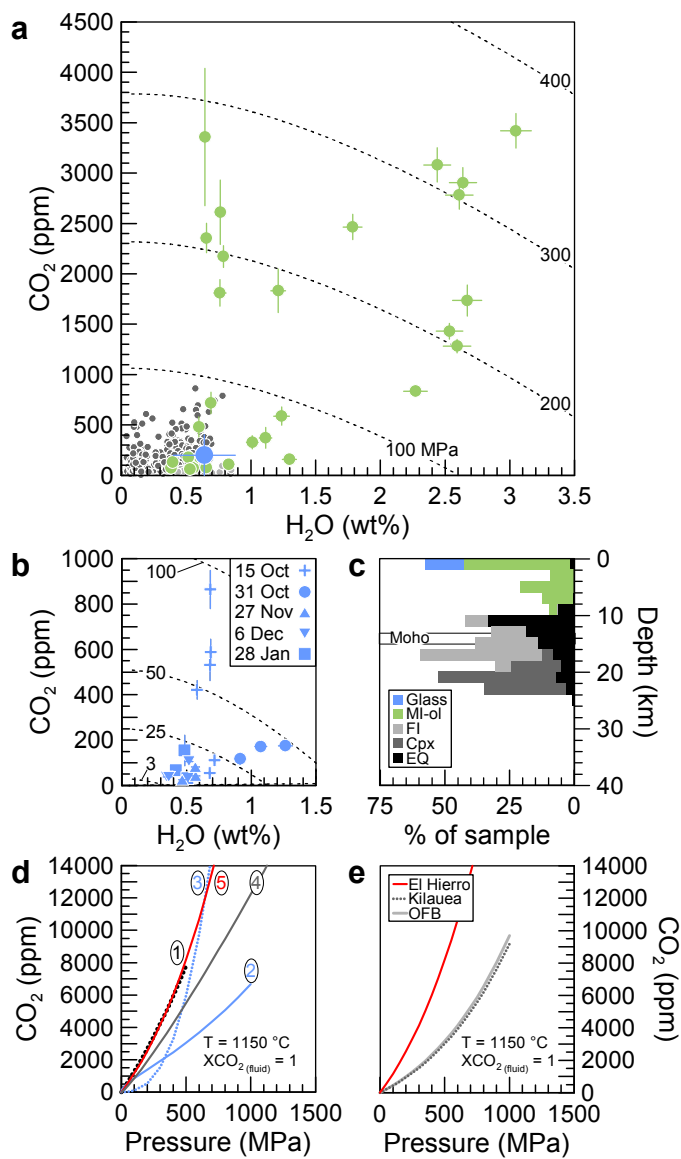
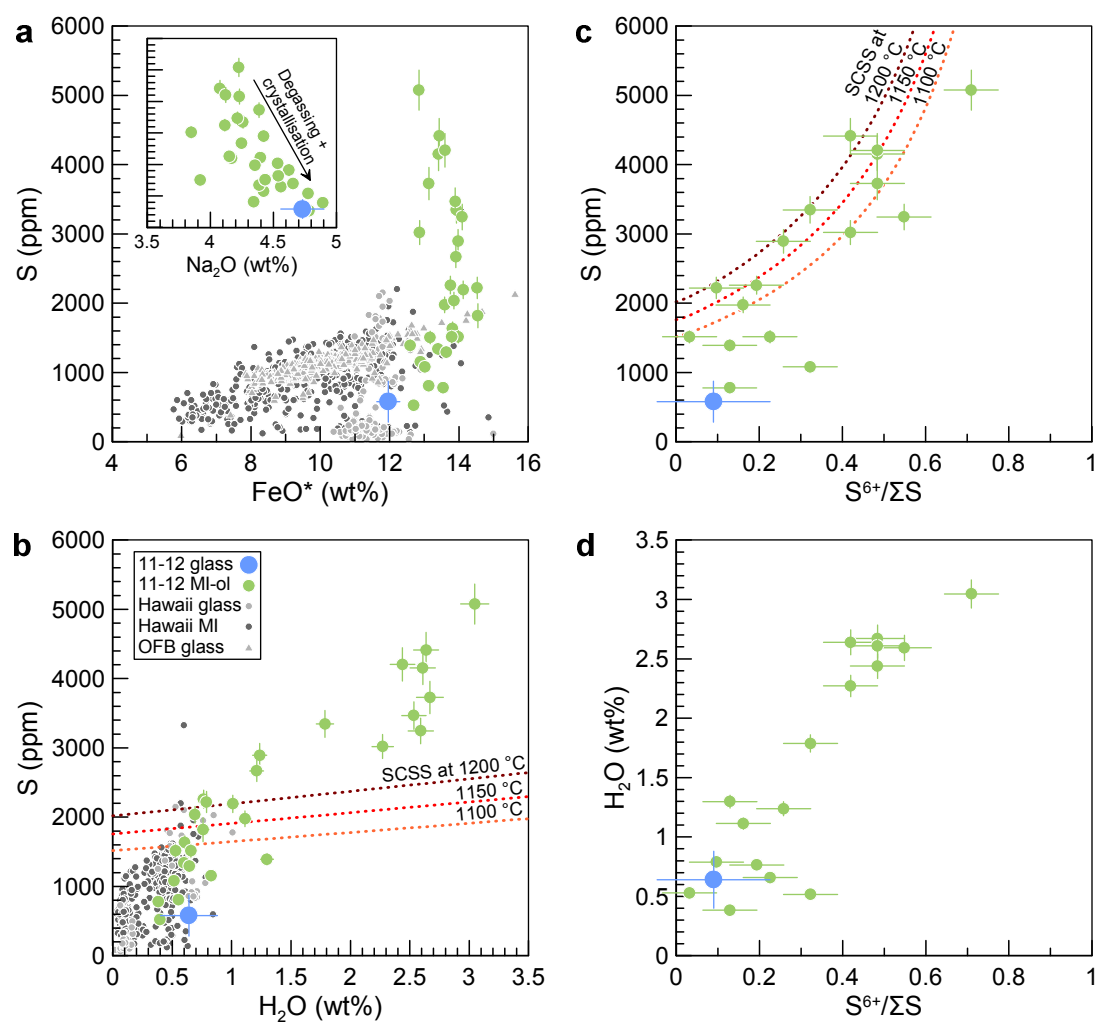


Figure 4

**Figure**[Click here to download Figure: Fig\\_5\\_Sulfur\\_revised\\_low.pdf](#)**Figure 5**

Figure

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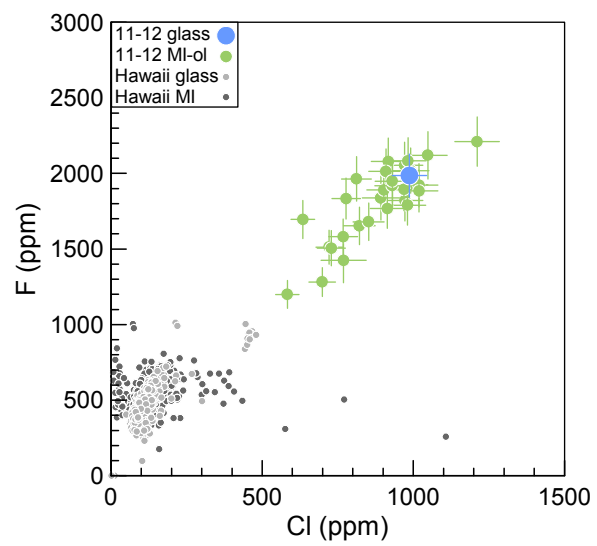
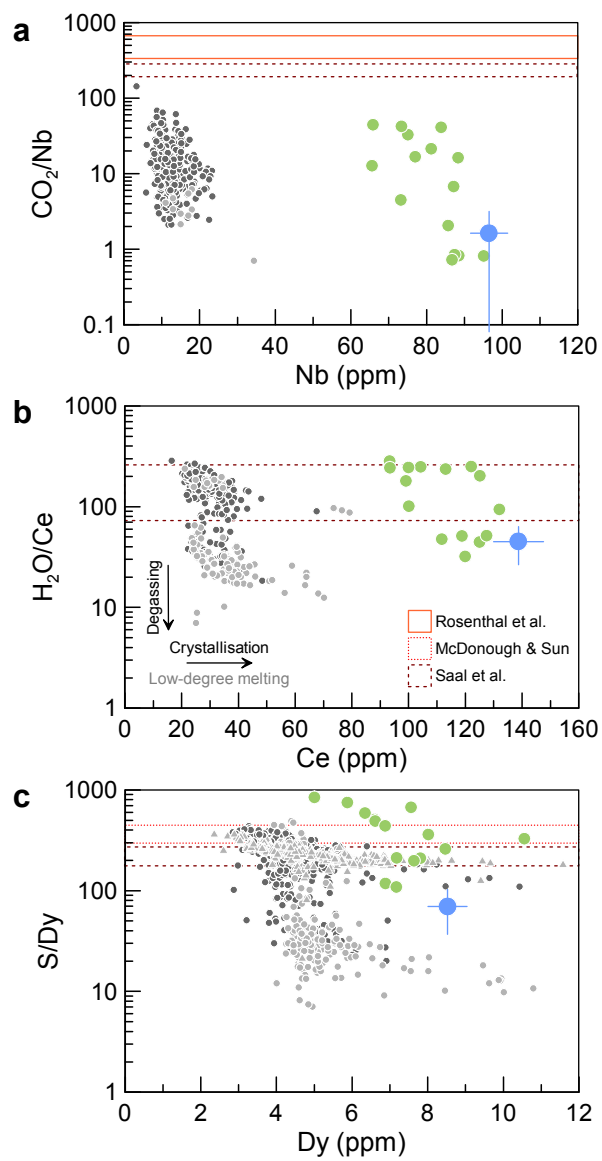


Figure 6

**Figure**

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**Figure 7**

Figure

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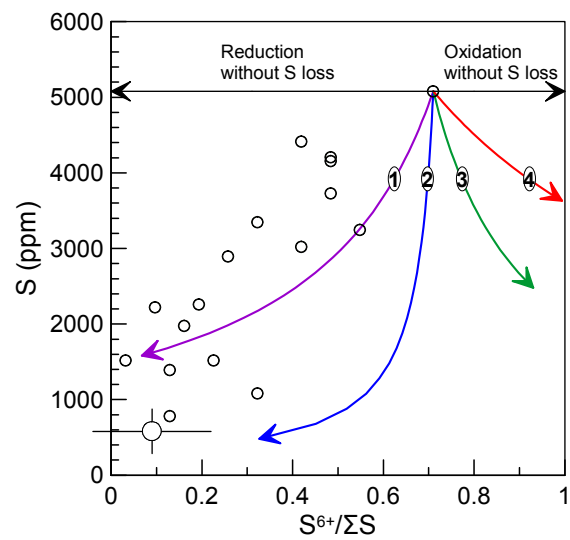


Figure 8

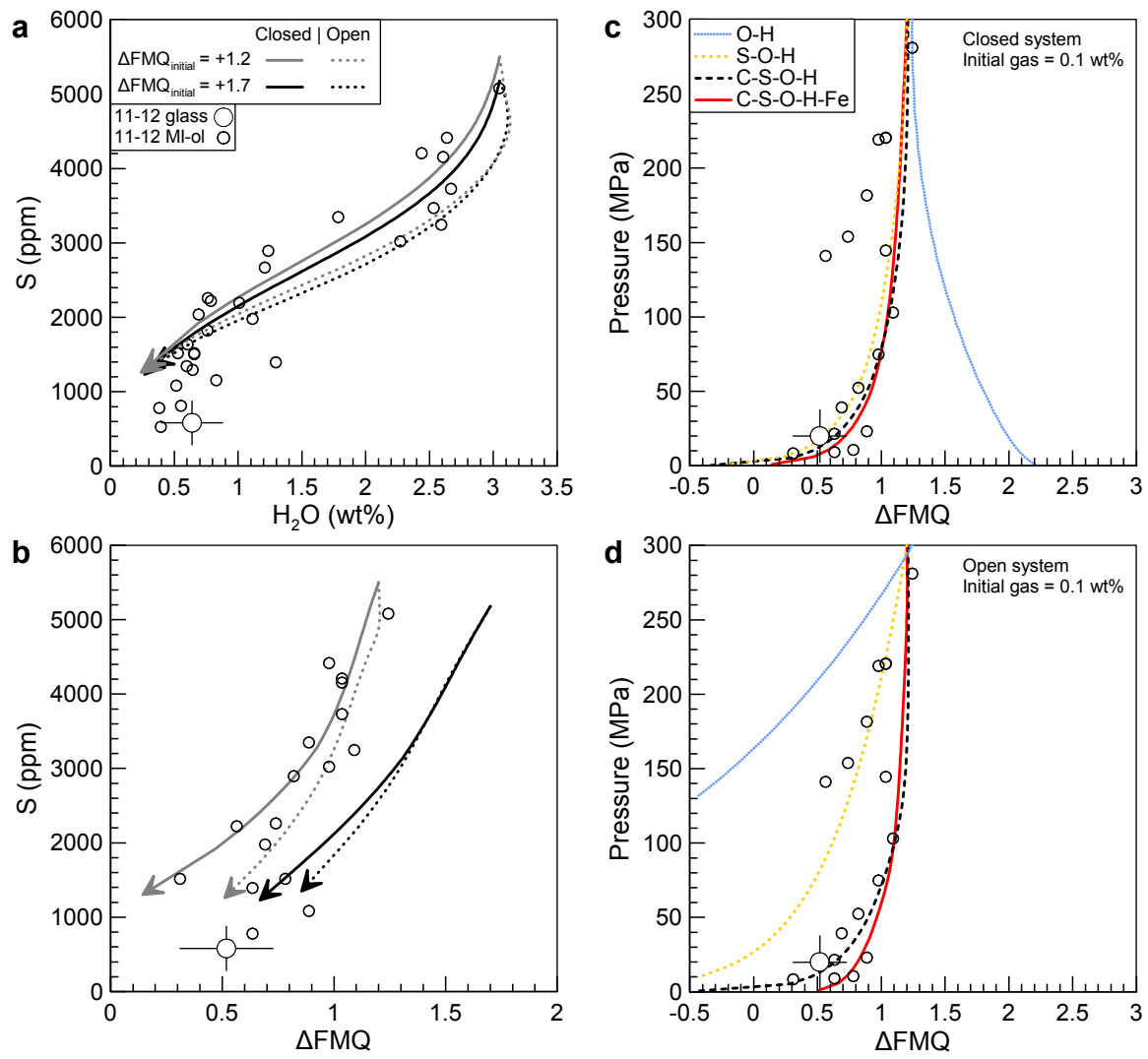


Figure 9



Figure

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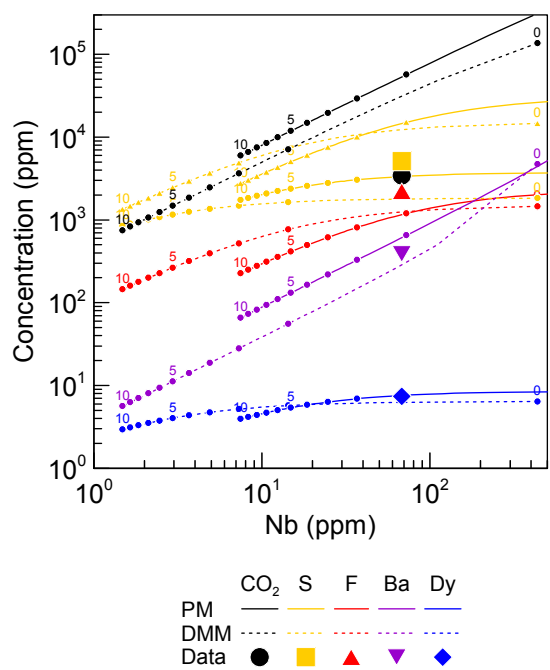


Figure 10